## **TFY4240**

## Solution problem set 3 Autumn 2015



## Problem 1.

a) We have a function f(x) on the interval [-1,1]. Since the Legendre polynomials form a complete set on this interval, it is possible to write f(x) as a linear combination of Legendre polynomials:

$$f(x) = \sum_{\ell=0}^{\infty} A_{\ell} P_{\ell}(x). \tag{1}$$

Multiplying by  $P_m$  on both sides and integrating from -1 to 1 gives

$$\int_{-1}^{1} dx f(x) P_m(x) = \int_{-1}^{1} dx \sum_{\ell=0}^{\infty} A_{\ell} P_{\ell}(x) P_m(x).$$
 (2)

Using the orthogonality of the Legendre polynomials, i.e.

$$\int_{-1}^{1} dx \, P_m(x) P_n(x) = \frac{2}{2n+1} \delta_{mn},\tag{3}$$

and interchanging the order of summation and integration in Eq. (2) leads to the relation

$$\int_{-1}^{1} dx f(x) P_m(x) = A_m \frac{2}{2m+1},\tag{4}$$

which is readily solved for the coefficients  $A_{\ell}$  to give (after renaming m to  $\ell$ )

$$A_{\ell} = \frac{2\ell + 1}{2} \int_{-1}^{1} dx \, f(x) P_{\ell}(x). \tag{5}$$

**b**)

$$f(x) = \begin{cases} -1 & x < 0 \\ +1 & x > 0 \end{cases} \tag{6}$$

Since f(x) is an odd function, and the Legendre polynomial  $P_{\ell}(x)$  is an even function for even  $\ell$ , the integrand in (5) is odd for even  $\ell$ , and thus  $A_{\ell}$  vanishes for this case. Hence only Legendre polynomials of odd order are needed, *i.e.* 

$$f(x) = \sum_{n=0}^{\infty} A_{2n+1} P_{2n+1}(x). \tag{7}$$

c) For odd  $\ell$ , the integral in equation (5) becomes twice the integral from 0 to 1, which gives  $A_{\ell} = (2\ell+1) \int_0^1 dx P_{\ell}(x)$ . The first few coefficients are

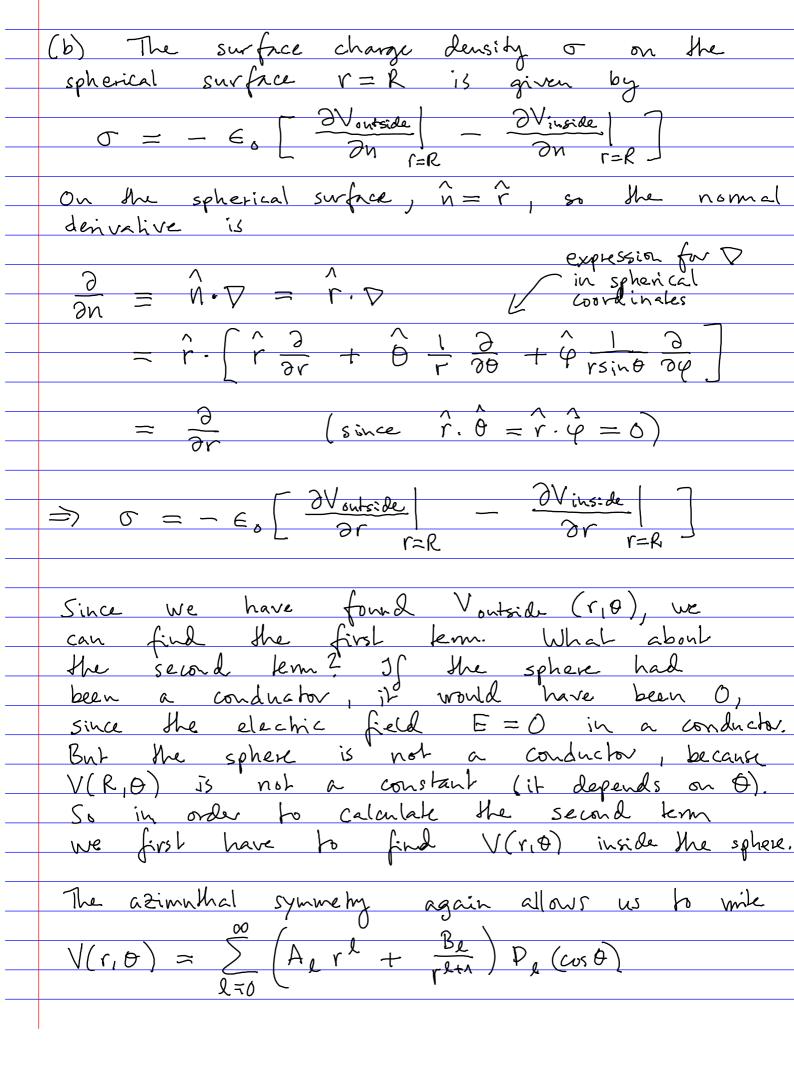
$$A_1 = 3 \int_0^1 dx \, x = \frac{3}{2},\tag{8}$$

$$A_3 = 7 \int_0^1 dx \, \frac{1}{2} (5x^3 - 3x) = \frac{7}{2} \left( \frac{5}{4} - \frac{3}{2} \right) = -\frac{7}{8}, \tag{9}$$

$$A_5 = 11 \int_0^1 dx \, \frac{1}{8} (63x^5 - 70x^3 + 15x) = \frac{11}{8} \left( \frac{63}{6} - \frac{70}{4} + \frac{15}{2} \right) = \frac{11}{16}. \tag{10}$$

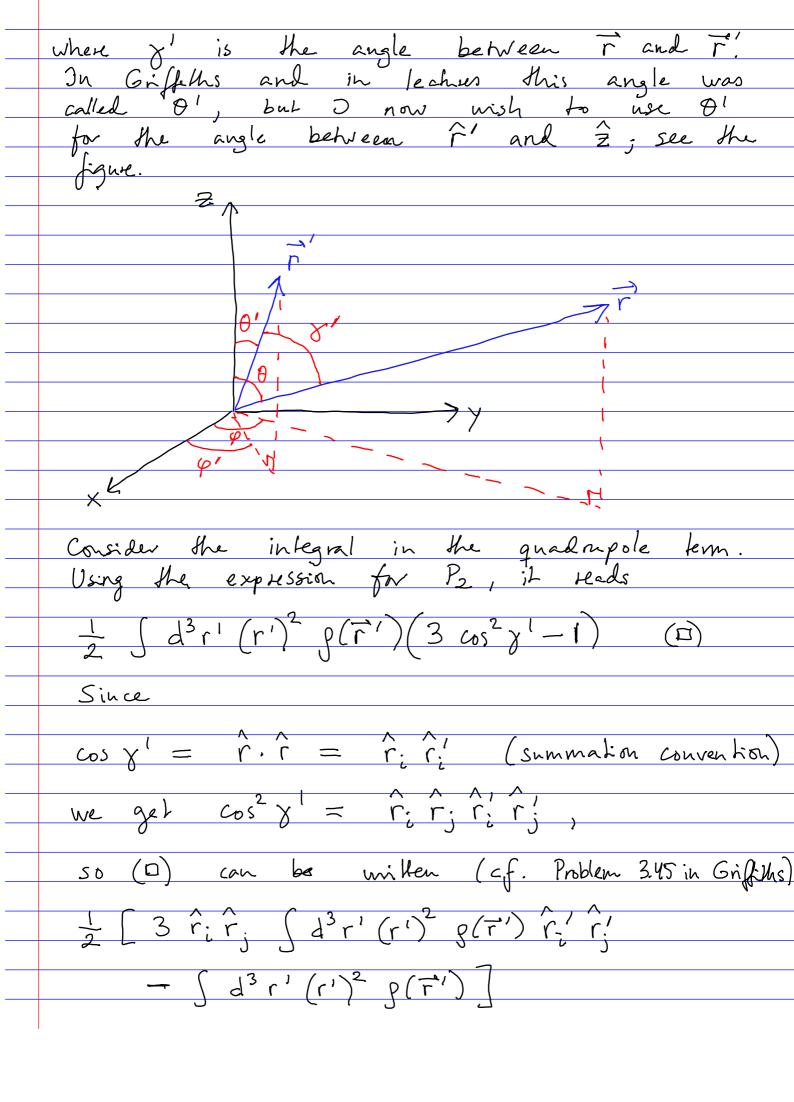
Problem 2 (a) Since the potential on the sphere surface is independent of the azimuthal angle  $\varphi$ , the problem has azimuthal symmetry. We can therefore expand the potential as  $V(r, \theta) = \sum_{l=0}^{\infty} \left( A_{l} r^{l} + \frac{B_{l}}{r^{l+1}} \right) P_{l} (\cos \theta)$ Since V(r,0) should  $\rightarrow 0$  as  $r \rightarrow \infty$ , At must be 0 for all l (including l=0)  $=) V(r,o) = \sum_{l=0}^{\infty} \frac{B_l}{\gamma^{l+1}} P_l(cos 0) \qquad (*)$ To find the coefficients Be, we need to consider the expansion (X) for  $\Gamma=R$  and use that it should equal  $V(R,D)=V_0\cos^2D$ . The coefficients could then be determined by using the orthogonality relations for Legendre polynomials. However, because in this example  $V(R,D)=V_0\cos^2D$  has a simple expression in terms of Legendre polynomials, we can use that to simply real off the coefficients. To this end, we note that  $P_o(\omega s \theta) = 1$ ,  $P_2(\omega s \theta) = \frac{1}{2}(3 \cos^2 \theta - 1)$  $\Rightarrow \cos^2\theta - \frac{2}{3}P_2(\cos\theta) = \frac{1}{3} = \frac{1}{3}P_0(\cos\theta)$  $\Rightarrow V(R_1\theta) = \frac{V_0}{3} \left[ P_0(cos\theta) + 2 P_2(cos\theta) \right]$ 

Comparing this with (x) evaluated at r=R, we can real off  $l=0 \text{ km}: \frac{B_0}{R} = \frac{V_0}{3} \implies B_0 = \frac{V_0 R}{3}$  $l=2 \text{ lem}: \frac{B_2}{R^3} = \frac{2V_0}{3} \implies B_2 = \frac{2V_0 R^3}{3}$  $l \neq 0, 2 : \frac{B_{\ell}}{\rho l + 1} \Rightarrow 0 \Rightarrow B_{\ell} = 0$ Inserting these results back into (x) gives  $V(r,\theta) = \frac{\sqrt{\sigma}}{3} \left[ \frac{R}{r} P_0(\cos \theta) + 2 \left( \frac{R}{r} \right)^3 P_2(\cos \theta) \right]$ Hence the potential outside the sphere (1>R) is Voutside  $(r, \theta) = \frac{\sqrt{\sigma}}{3} \left[ \frac{R}{r} + 2 \left( \frac{R}{r} \right)^{5} P_{2}(\cos \theta) \right]$ Let us check that this expression reduces to the correct result at r=R: Vousside  $(R_1\theta) = \frac{V_0}{2} \left[1 + 2P_2(\cos\theta)\right]$  $=\frac{\sqrt{0}}{3}\left[1+2\cdot\frac{1}{2}\left(3\cos^2\theta-1\right)\right]=\frac{\sqrt{0}\cos^2\theta}{0}$  ok



Be much be 0 for all l; otherwise  $V(r, \phi)$  would diverge as  $r \to 0$  $\Rightarrow V(r_1\theta) = \sum_{l=0}^{\infty} A_l r^l P_l(cos\theta) (**)$ Evaluating this at r=R and comparing with  $V(R_10) = \frac{V_0}{3} [P_0(\omega s 0) + 2P_2(\cos 0)]$  gives  $l=0 \text{ lem}: A_0 R^0 = \frac{V_0}{3} \implies A_0 = \frac{V_0}{3}$  $l = 2 \text{ term}: A_2 R^2 = \frac{V_0}{3} \cdot 2 \Rightarrow A_2 = \frac{2V_0}{3R^2}$  $l \neq 0,2 : A_{\ell} R^{\ell} = 0 \Rightarrow A_{\ell} = 0$ Inserting these results back into (XX) gives Vinside  $(r_1\theta) = \frac{\sqrt{o}}{3} \left[ P_0(\cos\theta) + 2\left(\frac{r}{R}\right)^2 P_2(\cos\theta) \right]$ (One can see that it reduces to Vo cos20 at r=R, as it should.) To summarize, we have found  $V_{\text{outside}}(V_1\theta) = \frac{V_0}{3} \left[ \frac{R}{r} + 2 \left( \frac{R}{r} \right)^3 P_2(\cos \theta) \right]$ Vinside  $(r_1\theta) = \frac{V_0}{3} \left[ 1 + 2\left(\frac{r}{R}\right)^2 P_2(\cos\theta) \right]$ This gives  $\frac{\partial V_{\text{outside}}}{\partial V} = \frac{V_0}{3} \left[ -\frac{R}{r^2} + 2 \cdot 3 \left( \frac{R}{r} \right) \cdot \left( -\frac{R}{r^2} \right) P_2(\cos \theta) \right]$  $\frac{\partial V_{\text{outside}}}{\partial V_{\text{r=R}}} = -\frac{V_0}{3R} \left[ 1 + 6 P_2(\cos \theta) \right]$ 

$$\frac{\partial V_{inside}}{\partial r} = \frac{V_r}{3} \cdot \frac{1}{2} \cdot \frac{1}{R} \cdot$$



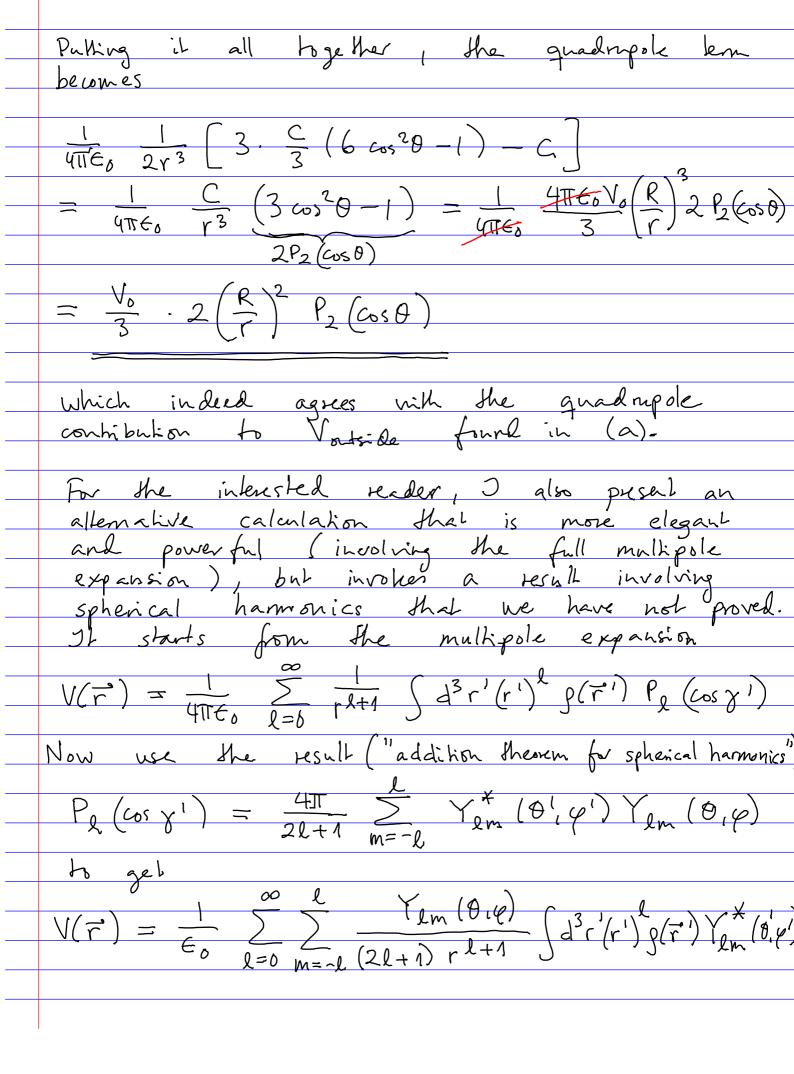
Note that by semiling in this way, all integrals are now independent of  $\overrightarrow{r}$ . The charge density is given by  $\rho(\vec{r}') = \sigma(\theta') \delta(r' - R).$  $\int d^{3}r' \left(r'\right)^{2} g\left(\overrightarrow{r'}\right) = \int dr'\left(r'\right)^{2} d\varphi' \sin \theta' d\theta' \left(r'\right) \delta\left(r'-R\right) \sigma(\theta')$  $\int_{Sde}^{e} = 2\pi R^{4} \int_{0}^{\pi} d\theta' \sin \theta' \sigma(\theta') = \frac{\epsilon_{0}V_{0}}{3R} \left[ P_{0}(\omega s\theta') + 10P_{2}(\omega s\theta') \right]$  $= 2\pi R^4 \frac{\epsilon_0 V_0}{3R} \int_{-\infty}^{\infty} d\theta' \sin\theta' \left[ P_0 \left( \cos\theta' \right) + 10 P_2 \left( \cos\theta' \right) \right]$  $=\frac{2\pi\epsilon_0 V_0 R^3}{3} \cdot \int dx \left[1 + 10 \frac{1}{2} \left(3x^2 - 1\right)\right]$  $= \frac{2\pi\epsilon_{0}V_{0}R^{3}}{3} \left[ 2 + 15\frac{1}{3}x^{3} \right] - 5 - 2 = \frac{4\pi\epsilon_{0}V_{0}R^{3}}{3}$ (As a check, note that this integral can (~ Vr) in Vouside. From this term we can tead off the total charge as 4TE, Vol., agreeing with our tesul above)

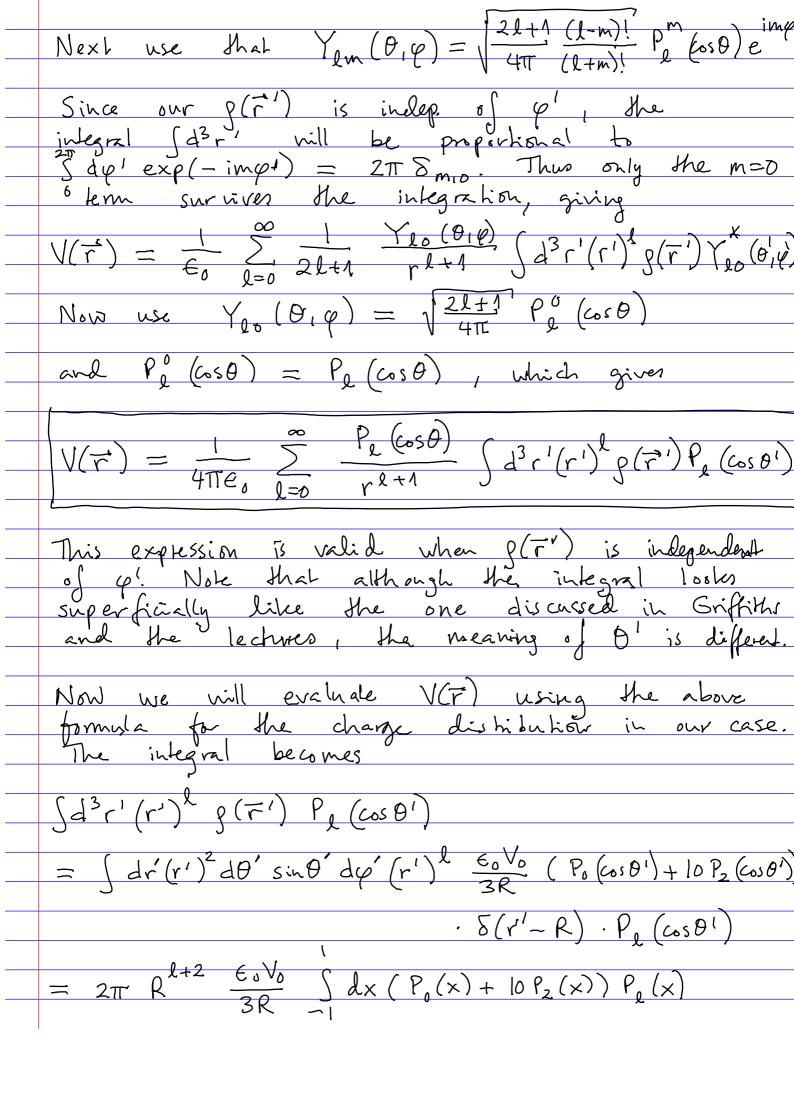
```
Next consider the sum
                          where \hat{\Gamma}_{x} = \sin\theta \cos\varphi, \hat{\Gamma}_{x}' = \sin\theta' \cos\varphi'

\hat{\Gamma}_{y} = \sin\theta \sin\varphi, \hat{\Gamma}_{y}' = \sin\theta' \sin\varphi'

\hat{\Gamma}_{z} = \cos\theta, \hat{\Gamma}_{z}' = \cos\theta'
                           There are 3.3 = 9 terms in the sum. Symbolically we can write the sum as (first letter i, second;)
         XX + yy + 22 + xy + yx
+ 2x + zy + x2 + y2
                                 We note that g(\vec{r}') is independent of \varphi'
= \sum_{i=1}^{2N} 2X_{i} \times 2X_{
                               Furthermore (define the shorthand K \equiv d^3r'(r')^2p(r'))
                               XX + yy = \frac{1}{\sin^2 \theta} \cos^2 \theta \int K \sin^2 \theta \int \cos^2 \theta' + \sin^2 \theta \int K \sin^2 \theta \int K \sin^2 \theta' \int \sin^2 
                            Since the average of \cos^2 \varphi' and \sin^2 \varphi' in [0,2\pi) is \frac{1}{2}, we can replace both factors by \frac{1}{2}
                    = ) xx + yy = \frac{1}{2} sin^2 \theta (cos^2 \varphi + sin^2 \varphi) \int K sin^2 \theta^{-1}
                                                                                                                                                                                                                 = \frac{1}{2} \sin^2 \theta \int K \sin^2 \theta'
```

```
= \frac{1}{2} \left( 1 - \cos^2 \theta \right) \int K \left( 1 - \cos^2 \theta' \right)
               Finally, ZZ = cos2 B S K cos2 b
                   So the sum reduces to
                        = (1-6520) JK(1-65201) + 6520 (K65201
            = \frac{1}{2} \left( \int K - \int K \cos^2 \theta' \right)
            + \cos^2\theta \left(-\frac{1}{2} \int K + \frac{3}{2} \int K \cos^2\theta'\right)
              We already calculated SK = \frac{4\pi\epsilon_0 V_0 R^3}{3} \equiv C so we just need to calculate
\int K \cos^2 \theta' = \int d^3 r' (r')^2 p(r') \cos^2 \theta'
= \ dr'(r')2 sin0 d0 dp (r)2 \ \frac{\xi_0\lambda_0}{20} \ \S(r'-R)
                                                                                                \cdot ( P_0 (\cos\theta) + 10 P_2 (\cos\theta)) \left(\frac{1}{3} P_0 (\cos\theta) + \frac{2}{3} P_2 (\cos\theta))
         = 2\pi \frac{\epsilon_0 V_0}{3R} \cdot R^{\frac{4}{1}} \frac{1}{3} \frac{
        = \frac{2\pi}{9} \in V_0 R^3 \left[ \frac{2}{2.0+1} + 10 \cdot 2 \cdot \frac{2}{2.2+1} \right] = \frac{2\pi}{3} \in V_0 R^3 \left[ \frac{2}{2.0+1} + 4 \right] = \frac{4\pi}{3} \in V_0 R^3 \left[ \frac{1}{3} + 4 \right] = \frac{4\pi}{3} = \frac{5}{3} \in V_0 R^3 \left[ \frac{1}{3} + 4 \right] = \frac{4\pi}{3} = \frac{5}{3} \in V_0 R^3 \left[ \frac{1}{3} + \frac{4}{3} + \frac{4}{3} + \frac{5}{3} + \frac{5}{3}
        Thus the sum becomes
              \frac{1}{2} \zeta \left( 1 - \frac{5}{3} \right) + \frac{1}{2} \cos^2 \theta \left( -C + 3 \cdot \frac{5}{3} C \right) = \frac{C}{3} \left( 6 \cos^2 \theta - 1 \right)
```

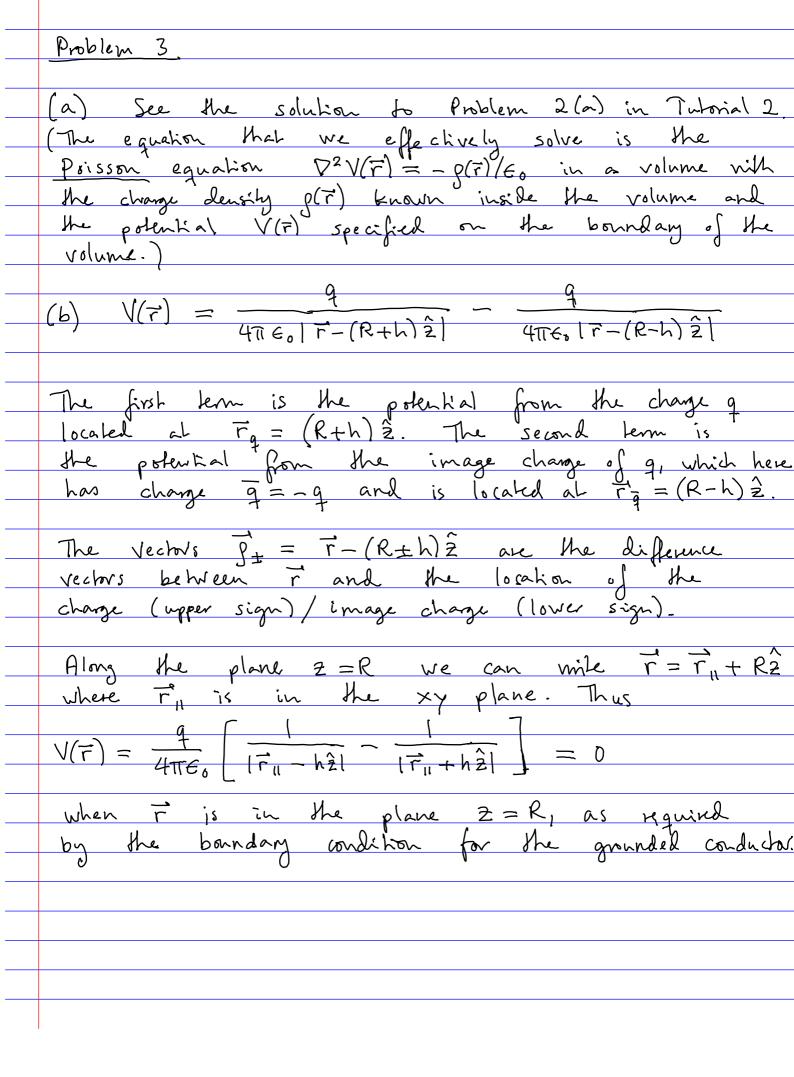


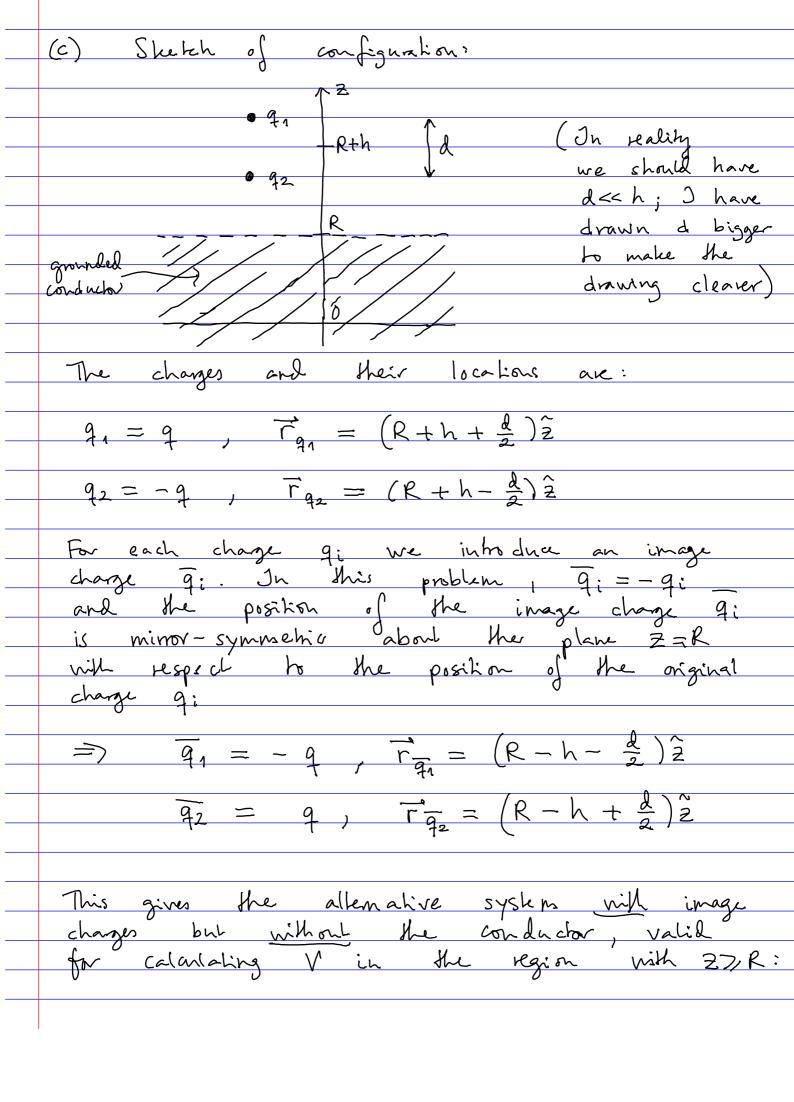


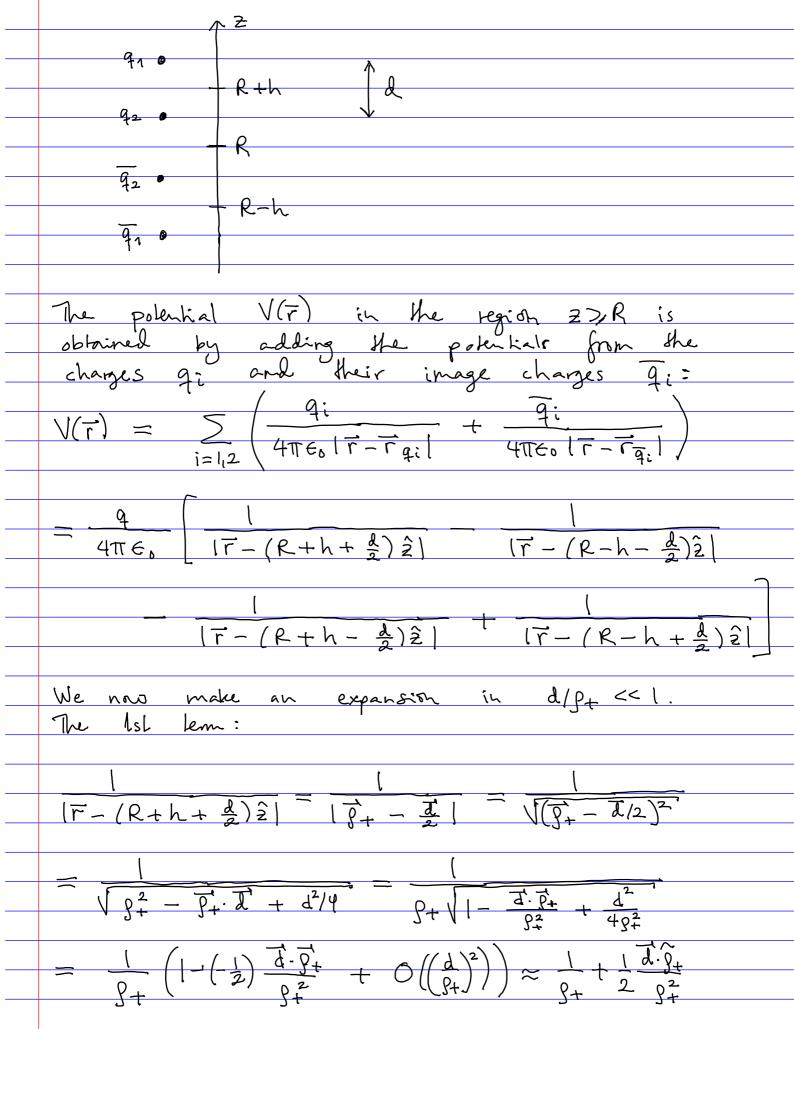
$$= 2\pi \epsilon_{0} V_{0} R^{l+1} \left[ \frac{2}{2 \cdot 0 + 1} \delta_{l,0} + 10 \cdot \frac{2}{2 \cdot 2 + 1} \delta_{l,2} \right]$$

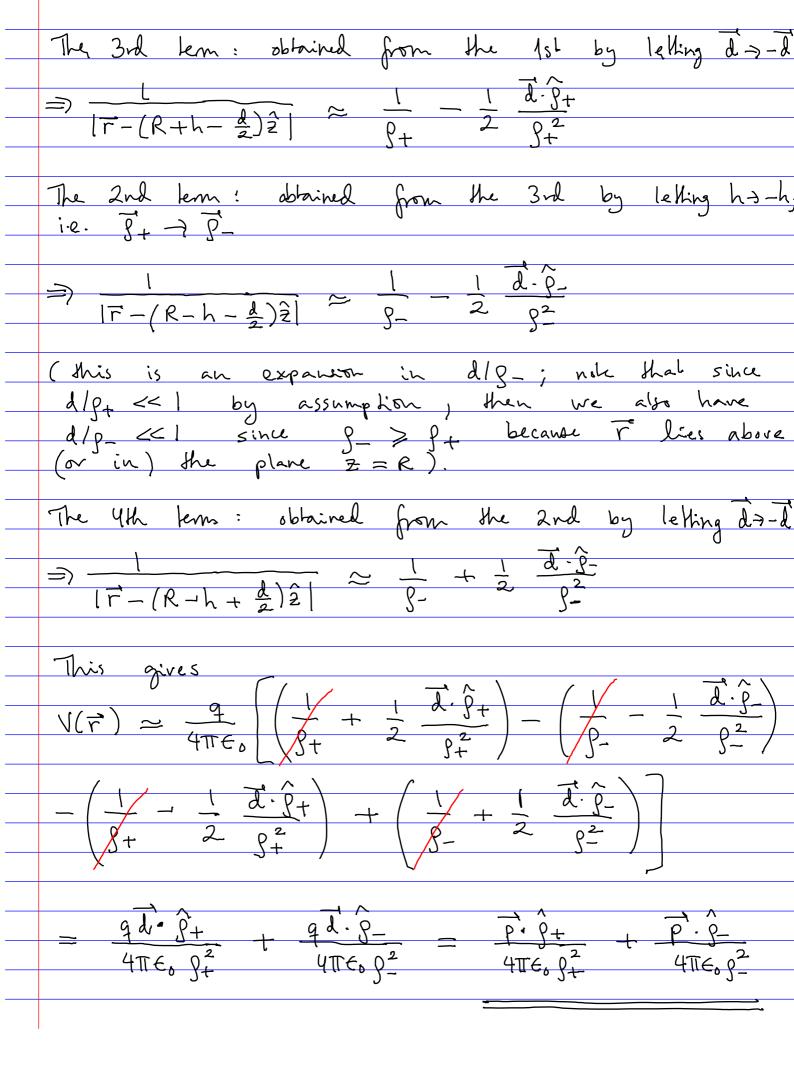
$$= \frac{4\pi \epsilon_{0} V_{0} R^{l+1}}{3} \left( \delta_{l,0} + 2 \delta_{l,2} \right)$$
This gives
$$V(\vec{r}) = \frac{1}{4\pi \epsilon_{0}} \sum_{l=0}^{\infty} \frac{P_{l}(cos\theta)}{r^{l+1}} \cdot \frac{4\pi \epsilon_{0} V_{0} R^{l+1}}{3} \left( \delta_{l,0} + 2 \delta_{l,2} \right)$$

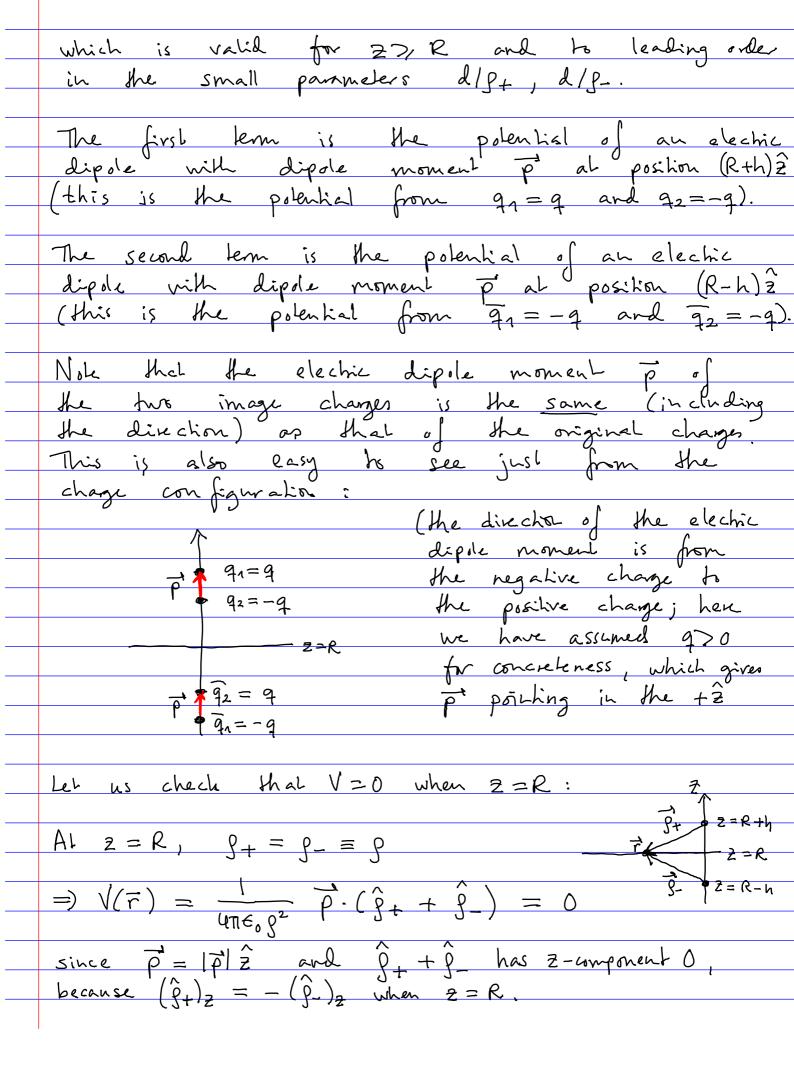
$$= \frac{V_{0}}{3} \left[ \frac{R}{r} + 2 \left( \frac{R}{r} \right)^{3} P_{2}(cos\theta) \right]$$
which agrees with our result for Vankide in (a).











(d) By symmetry, the image charge 
$$\bar{q}$$
 must be positioned on the  $Z$  axis. We can therefore trick its position as  $\bar{r}_q = Z_q \hat{Z}$ , where  $|Z_q| < R$  since the image charge must be outside the region of interest (the outside of the sphere). Also let us write  $\bar{r}_q = Z_q \hat{Z}$  where  $Z_q = R + h$ . The trial posterial outside the sphere can then be written as 
$$V(\vec{r}) = \frac{q}{4\pi\epsilon_0} \frac{q}{|r\hat{r} - Z_q \hat{Z}|} + \frac{q}{|r\hat{r} - Z_q \hat{Z}|} = \frac{1}{|r\hat{r} - Z_q \hat{Z}|} = \frac{1}{|r\hat$$

$$\frac{1}{2} = \frac{R^2}{Z_q} = \frac{R^2}{R+h} = \frac{1}{R+h/R}$$

$$\frac{1}{q} = -\frac{1}{R} = \frac{1}{q} = -\frac{R}{R+h} = \frac{1}{1+h/R} = \frac{1}{1+h/R} = \frac{1}{2} = \frac{1}{R} = \frac{1}{1+h/R} = \frac{1}{1+h/R} = \frac{1}{R} = \frac{1}{R} = \frac{1}{R} = \frac{1}{R+h/R} = \frac{1}{1+h/R} = \frac{1}{R} = \frac{1}{R+h/R} = \frac{1}{1+h/R} = \frac{1}{1+h/R} = \frac{1}{1+h/R} = \frac{1}{1+h/R} = \frac{1}{R+h/R} = \frac{1}{1+h/R} = \frac{1}{1+h/R}$$

This corresponds to also restricting it to positions  $\vec{\Gamma} = \vec{\Gamma}_{11} + (R + \Delta z)^{2} \qquad \text{with} \quad \Delta z > 0$ and then letting R & while TI, DZ and he are kept fixed. Let us check that in this limit the potential outside the sphere reduces to the potential above the infinite flat conducting plane: T-(R+h)== T,+(R+Dz)=-(R+h)== T,+(Dz-h)= (independent of R, and thus constant as  $R \rightarrow \infty$ ),  $\vec{r} - R \frac{1}{1 + h/R} \hat{z} = \vec{r}_{11} + (R + bz) \hat{z} - R \frac{1}{1 + h/R} \hat{z}$  $\approx \vec{r}_{\parallel} + (R + \Delta z)^{2} - R(1 - \frac{h}{R})^{2} \approx R \rightarrow \infty$  $= \vec{\Gamma}_{\parallel} + (\Delta z + h) \hat{z} \qquad \approx R \rightarrow \infty$ Also, the factor  $\frac{1}{(1+h/R)}$  in the image charge  $\rightarrow 1$  as  $R \rightarrow \infty$  $\Rightarrow \sqrt{(r)} \xrightarrow{R \to \infty} \underbrace{q}_{4\pi\epsilon_0} \left[ \frac{1}{[\vec{r}_{11} + (\Delta z - h)\hat{2}]} - \frac{1}{[\vec{r}_{11} + (\Delta z + h)\hat{2}]} \right]$ which indeed agrees with the expression for  $V(\vec{r})$  found for the system studied in (b).

(f) (It is implicit that both charges are outside the sphere, i.e. d/2 < h.) We can use the usults found in (d) to obtain V(F) for this case too. For each charge  $q_i$ , at position  $\overrightarrow{r}_{q_i} = Z_q$ ,  $\widehat{Z}$ , we have an image charge q; at position r= = = = = = , where  $\overline{q}_i = -\frac{R}{2q_i} q_i$ ,  $Z_{\overline{q}_i} = \frac{R^2}{2q_i}$ The potential V(r) outside the sphere is then  $V(\vec{r}) = \sum_{i=l,2} \frac{q_i}{4\pi\epsilon_o |\vec{r} - \vec{r}_{q_i}|} + \frac{\overline{q}_i}{4\pi\epsilon_o |\vec{r} - \vec{r}_{\overline{q}_i}|} (x)$ The monopole term in the multipole expansion for  $V(\vec{r})$  is  $Q/477\epsilon_0 r$  where Q is the total charge due to all charges in (\*), i.e. Q = 91 + 92 + 91 + 92 Using  $q_1 = q_1 + q_2 = -q_1$  $\frac{1}{q_1} = -\frac{R}{2q_1} q_1 = -\frac{R}{R+h+d/2} q_1$  $q_2 = -\frac{R}{2g_2} q_2 = \frac{R}{R + h - d/2} q_1$  $Q = \overline{q}_1 + \overline{q}_2 = qR\left(\frac{1}{R+h-d/2} - \frac{1}{R+h+d/2}\right)$ Since the expussion inside the parenthesis is >0, Q has the same sign as q, i.e. positive.

Q is the total change of the system. Since the system consists of the changes  $g_1$  and  $g_2$  and the conducting sphere, it follows that Q = 91 + 92 + 95 where  $q_s$  is the charge induced on the surface of the conducting sphere. Since  $q_2 = -q_1$ ,  $\frac{Q = q_s}{2}$ This is the physical interpretation of Q. We found Q > 0. This can be qualitatively under stood as follows: The two changes quark go will each induce surface changes on the sphere with opposite sign of gi. And the closer to the sphere the charge is, the bigger the magnitude of the induced surface charge. Since the negative charge go is closer to the sphere than the positive charge go, the net induced surface charge 9s is positive.